## PhyzExamples: UCM & Gravity

### Physical Quantities • Symbols • Units • Brief Definitions

**Radius** • R or  $r \cdot m \cdot D$ istance from the center of the circle to the center of the object in circular motion. If motion is circular, radius is constant.

**Tangential Velocity** • v / **Tangential Speed** •  $v \cdot m/s$  • Velocity or speed of an object in circular motion. In uniform circular motion, tangential *speed* is constant while tangential *velocity* is always changing.

**Centripetal Force** •  $F_C$  • N • Force on an object in circular motion directed radially inward (toward the center of the circle).

**Gravitational Force** •  $F_G$  or  $W \cdot N \cdot Attractive force between two objects with mass (due only to their mass).$ 

Universal Gravitation Constant •  $G \cdot N \cdot m^2/kg^2 \cdot 6.67 \times 10^{-11} N \cdot m^2/kg^2$ 

### Equations

 $F_C = mv^2/R$  • centripetal force = mass · tangential speed squared / radius

 $F_G = GMm/R^2 \bullet$  (Newton's Universal Gravitation) • gravitational force = universal gravitation constant • mass of one object • mass of the other object / radius squared

### Smooth Operations Examples

1. What is the force needed to keep a 0.50-kg object moving in a 0.60-m circle at 2.8 m/s?

1. m = 0.50 kg R = 0.60 m v = 2.8 m/s  $F_c$  = ?  $F_c$  = mv<sup>2</sup>/R  $F_c$  = 0.50 kg · (2.8 m/s)<sup>2</sup> / 0.60 m  $F_c$  = 6.5 N 2. What is the mass of an object moving at 4 m/s in a circle having a 3-m radius and experiencing a 24-N centripetal force?

2. v = 4 m/s R = 3 m F<sub>c</sub> = 24 N m = ? F<sub>c</sub> =  $mv^2/R$ m = F<sub>c</sub> R/ $v^2$ m = 24 N · 3 m / (4 m/s)<sup>2</sup> <u>m = 4.5 kg</u>

4. How far is 80-kg Felix from 60-kg Bertha if they experience an attractive gravitational force of 100 nN?

4.  $M = 80 \text{ kg} \text{ m} = 60 \text{ kg} \text{ F}_G = 100 \times 10^{-9} \text{ N}$  R = ?  $F_G = GMm/R^2$   $R = \sqrt{(GMm/F_G)}$   $R = \sqrt{(6.67 \times 10^{-11} \text{ N} \cdot \text{m}^2/\text{kg}^2 \cdot 80 \text{ kg} \cdot 60 \text{ kg} / 100 \times 10^{-9} \text{ N})}$ R = 1.79 m

3. What is the gravitational force of the Earth on the moon?

[Numerical values found on Solar System Reference sheet]

3. M = 5.98 x  $10^{24}$  kg m = 7.36 x  $10^{22}$  kg R = 3.82 x  $10^{8}$  m F<sub>G</sub> = GMm/R<sup>2</sup> F<sub>G</sub> = 6.67 x  $10^{-11}$  N·m<sup>2</sup>/kg<sup>2</sup> · 5.98 x  $10^{24}$  kg · 7.36 x  $10^{22}$  kg /  $(3.82 \times 10^{8} \text{ m})^{2}$ F<sub>G</sub> = 2.0 x  $10^{20}$  N

OF	<b>RBITAL CHAI</b>	ORBITAL CHARACTERISTICS	S		PLA	NETARY CH/	PLANETARY CHARACTERISTICS	ICS
Radius	ius	Period	Speed		Mass	Radius	Density	Gravity
-9m*	AU	years	m/s	E	E+24kg*	$E+6m^*$	kg/m^3	m/s^2
7.9	0.39	0.241	47,900		0.334	2.44	5,600	
108	0.72	0.615	35,000		4.87	6.1	5,200	
150	1.00	1.00	29,800		5.98	6.4	5,520	9.78
228	1.52	1.88	24,100		0.640	3.4	3,950	
178	5.19	11.9	13,100		1902	72	1,310	
430	9.53	30	9,640		569	60	704	
870	19.1	84	6,810		86.7	25.9	1,210	
500	30.0	165	5,430		103	24.8	1,670	
				[				

## ODBITAL CHADACTEDICTICS

E+S

S

**PhyzReference: Solar System Information** 

*Note: "E" means "times 10 to the," so	the," so
57.9E+9m is 57.9 times 10 to the 9th meters.	we 9th meters.

45

28

14

Moon (Luna) Sun (Sol)

the planets, the Sun, and the Moon. Assignment: Calculate gravity for

## Earth-Moon Distance: 3.82E+8m

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# SOL & LUNA: THE SUN AND THE MOON

Mass	Radius	Density	Gravity
kg	т	kg/m^3	m/s^2
1.99E+30	6.96E+8	1,410	
7.36E+22	1.74E+6	3,340	

Mercury Neptune Uranus Saturn Jupiter Venus Earth Mars

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